IYA 2009 "You are Galileo Project"

Let's Observe M31 and M45!

Name

Age

Address

Observation and Sketch

An 18 century French Astronomer, Charles Messier made a	
catalogue of celestial objects that appeared blurry. These	
objects were later called by their number with a "M" at the begin-	
ning. Later astronomers have found that his catalogues included	
many different kinds of celestial objects. The true nature of M31	
was especially ground-breaking. Let's observe and compare	
M31 and M45 and try to discover their true natures.	

Month	Day	:	M45	Date Site	Month	Day	:
cope	cm_Power	<u>x</u>	Diamete	r of Telesco	ope	Power	х
		Month Day	Month Day :	Month Day : M45	Month Day : Date M45 Site Site Diameter of Telesco	Month Day Date Month	Month Day : Date Month Day

* The Power of the Telescope can be calculated by dividing the focal length of the telescope with the focal length of the eyepiece.

* M31 is darker then M45, so it's hard to find and see.

If you couldn't see it this time, try again when you are visiting places that are darker in the night.

■Write down what you noticed, and anything peculiar you found.



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Let's Observe M31 and M45!

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■The photoes below are of M31 and M45 taken using a telescope with a large diameter.

Center of M31



Photoes provided by National Astronomical Observatory of Japan

■Let's list the characteristics of M31 and M45 from our sketches, observations and photoes above.

M31

M45



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■M45

M45 is also known as Pleades, or "Seven Sisters" (in English), and "Subaru" (in Japanese) You can see that M45 is a cluster of stars. It includes a couple hundred stars and it resides at around 400 light years* from Earth.

How many stars could you count with your Telescope?

(*1 light year = the distance light travels in a year = 9.5 trillion km = 5.9 trillion miles)

∎M31

M31 looks spread out and blurry. You cannot tell whether it is a cluster or not from the pictures or sketches, as you could with M45.

Observations performed in the 19th century proved that M31 was also a cluster of stars.

The reason M31 was blurry and the stars could not be distinguished was because M31 was even further than M45.



There was a debate held between two American astronomers, Harlow Shapley and Herber Curtis, over how far M31 was. Did M31 reside within our Milky-Way Galaxy, or did it exists outside of our galaxy independently?



■ This problem was later solved by Hubble. He used the two following facts.

1. Cephide variable stars

These are stars that turn bright and then dark cyclically. The length of the cycle and brightness correlate: brighter stars have a longer cycle. This means we can calculate the true brightness of a star from its cycle.



2. Relationship between apparent brightness and distance.

As the figure below shows, light disburses as distance increase, If the distance doubles, the brightness decreases to 1/4. If the distance triples, the brightness decreases to 1/9. This means that the apparent brightness of a star is inversely proportional to the square of distance. Thus, we can calculate the distance by comparing the actual brightness of a star and its apparent brightness.





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(Earth)

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with a diameter of 7.6cm)

Let's Observe M31 and M45!

↓ An Overview of M31 (photographed using a telescope

You can see that it's center is very bright.

Around 120,000 light years

Image provided by National Astronomical Observatory of Japan.

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What is the true nature of M31 that is so large and far away? Let's consider this by comparing M31 against our Milky-Way Galaxy.



What is M31? Let's write down our thoughts.

This is a image of a part of M31 photographed using the Subaru Telescope.

You can see individual stars as white dots!



■Write down what you have learned, and what you want to learn in the future.

